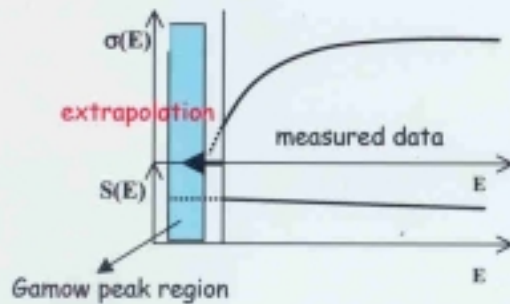
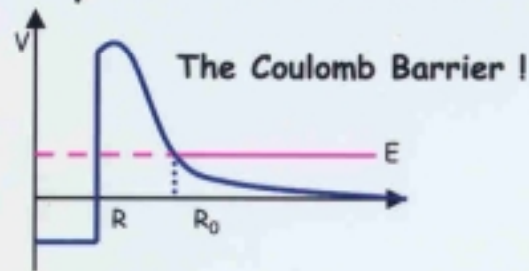


Determination of Nuclear Cross Section at Astrophysical Energies

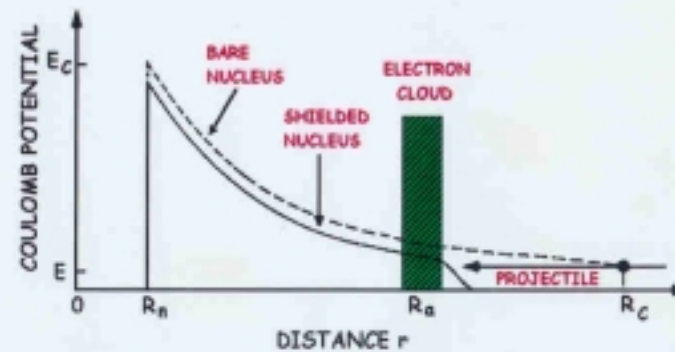
We have to face several problems

Very low Cross Section



$$\sigma(E) = \pi \lambda^2 \cdot P \cdot S(E)$$

Presence of the Electron Screening Effect



Result: Enhancement of the Astrophysical Factor

$$f(E) = \frac{S_s(E)}{S_b(E)} = \exp\left(\pi\eta \frac{U_e}{E}\right)$$

The Trojan-Horse Method

G. Baur *Phys. Lett. B* 78, 35(1986)

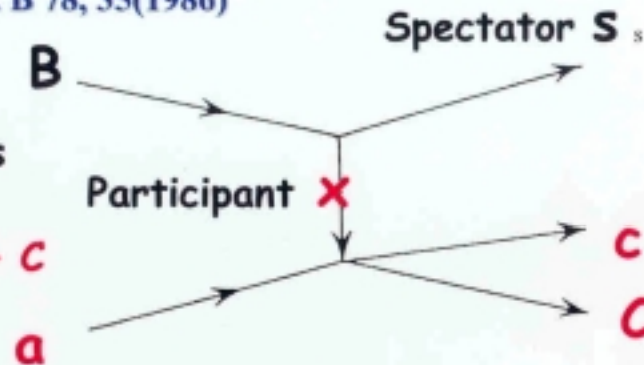
- Quasi-free mechanism**

three-body reaction $a + B \rightarrow c + C + s$

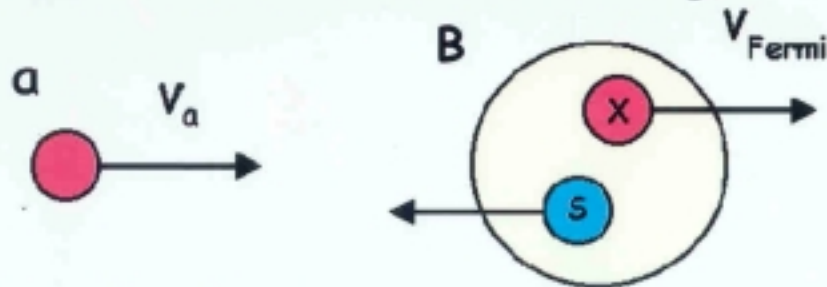
(with B composed by two clusters $x \oplus s$)

in order to study the 2-body $a + x \rightarrow c + C$

reaction at the astrophysical energies



- If $E_a > E_{Coul} \Rightarrow$ we can neglect Coulomb Effects...
(Coulomb Barrier and Electron Screening)



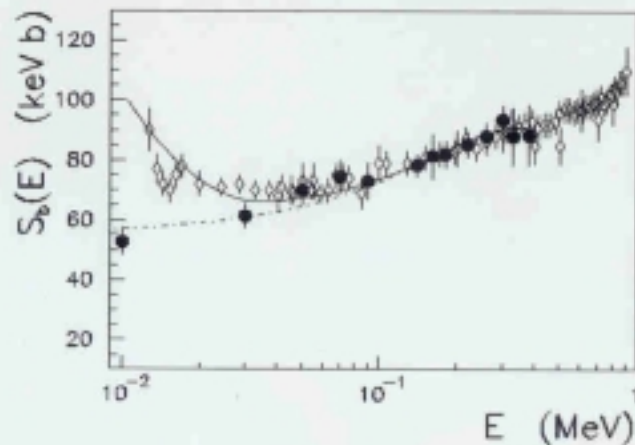
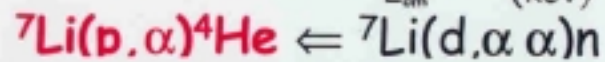
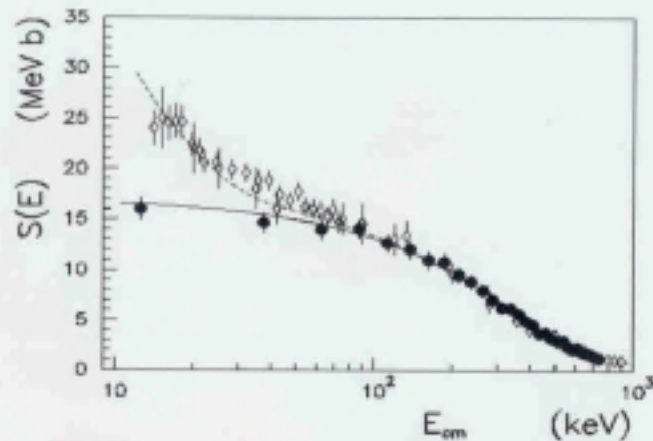
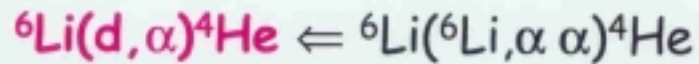
- but we can have $E_{ax} \approx 0 \Rightarrow$ due to the internal Fermi Motion

The Link between the two-body and the three-body (TH) reaction

$$\underbrace{\frac{d^3\sigma}{d\Omega_c d\Omega_s dE_c}}_{\text{Measured}} \propto \underbrace{(KF) \cdot |G(\vec{P}_s)|^2}_{\text{Calculated}} \cdot \underbrace{\left(\frac{d\sigma}{d\Omega}\right)_{x(a,c)C}}_{\text{The two-body relevant Cross-Section}}$$

KF= Kinematic Factor

$|G(\vec{P}_s)|^2$ = Momentum Distribution of s inside B



$${}^6\text{Li} = d \oplus \alpha$$

$$U_e = 320 \pm 51 \text{ eV}$$

$$U_{\text{theo}} = 186 \text{ eV}$$

$$S(0) = 16.9 \pm 0.5 \text{ MeV} \cdot \text{b}$$

- ◊ (Engstler S. et al.: 1992, Z. Phys., A342, 471)
- (C. Spitaleri et al.: 2001, new results)

$$d = p \oplus n$$

$$U_e = 330 \pm 40 \text{ eV}$$

$$U_{\text{theo}} = 186 \text{ eV}$$

$$S_0 = 55 \pm 3 \text{ keV} \cdot \text{b}$$

- ◊ (Engstler S. et al.: 1992, Z. Phys., A342, 471)
- (Aliotta M. et al.: 2000, Eur.Ph.J. 9, 435)

Summary

U_e (theo) (ad. Limit).	U_e ${}^6\text{Li}+d$ (THM)	U_e ${}^6\text{Li}+d$ (DIRECT)	$S(0)$ ${}^6\text{Li}+d$ (THM)	$S(0)$ ${}^6\text{Li}+d$ (DIRECT)
186 eV	320 ± 50 eV	380 ± 250 eV	16.9 ± 0.5 MeV b	17.4 MeV b

- We confirm the extrapolation trend for $S(E)$;
- We find a good agreement between $U_{e(\text{THM})}$ and $U_{e(\text{DIRECT})}$ previously measured (but with a smaller error bar for $U_{e(\text{THM})}$);
- We find again discrepancies between experimental e theoretical determinations of U_e (troubles with the atomic models ???);
- we find again no isotopic dependence of the Screening Effect.

Conclusions

- 😊 Reduced Coulomb barrier effects
- 😊 Improved information on Electron Screening
- 😞 Normalization to Direct Data
- 😞 Discrimination from "Background" Reactions

Perspectives...



The reaction ${}^6\text{Li}(p, \alpha){}^3\text{He}$ (depletion ${}^6\text{Li}$ e screening);

The reaction ${}^3\text{He}(d, p){}^4\text{He}$ (screening).