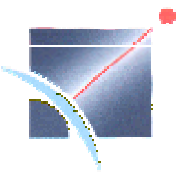


Bose-Einstein condensation of metastable He



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EPS-12 August 2002

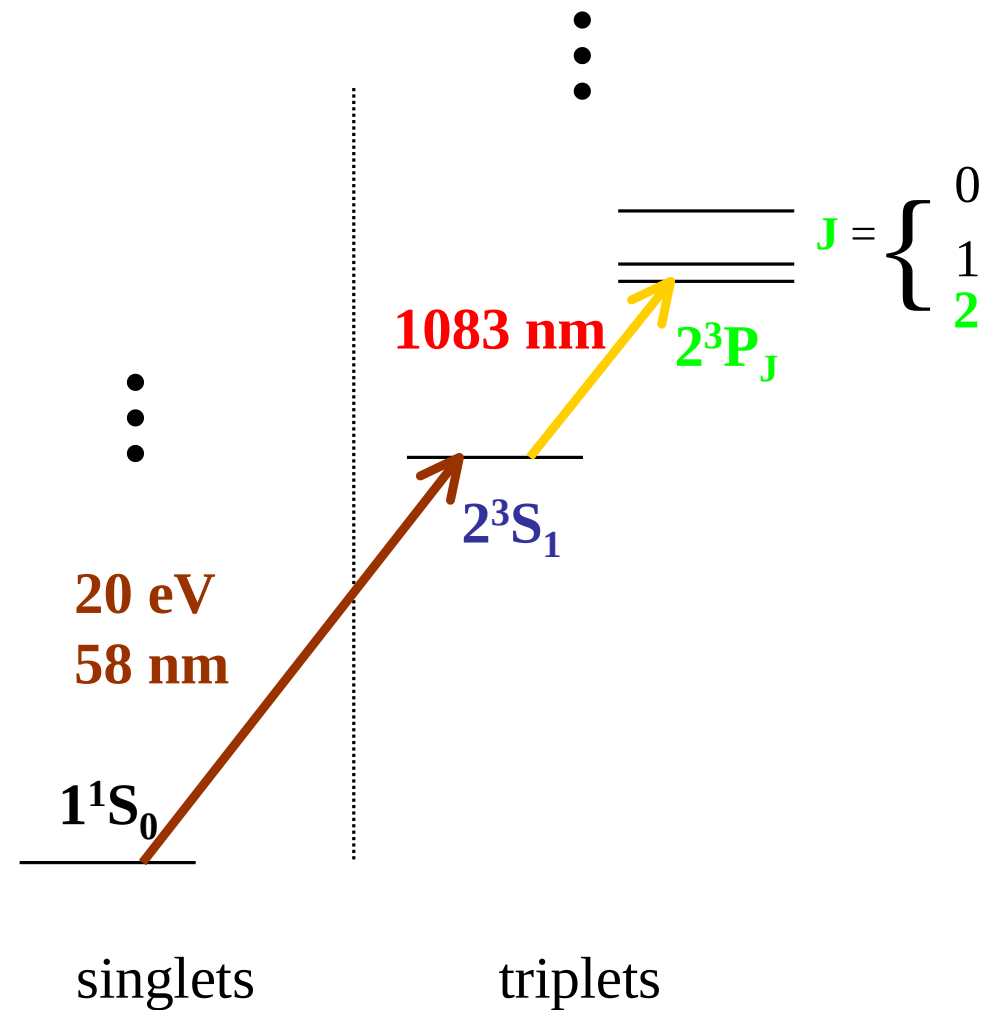
the helium atom

lifetimes:

$$2^3S_1 \rightarrow 1^1S_0 : \sim 9000 \text{ s}$$

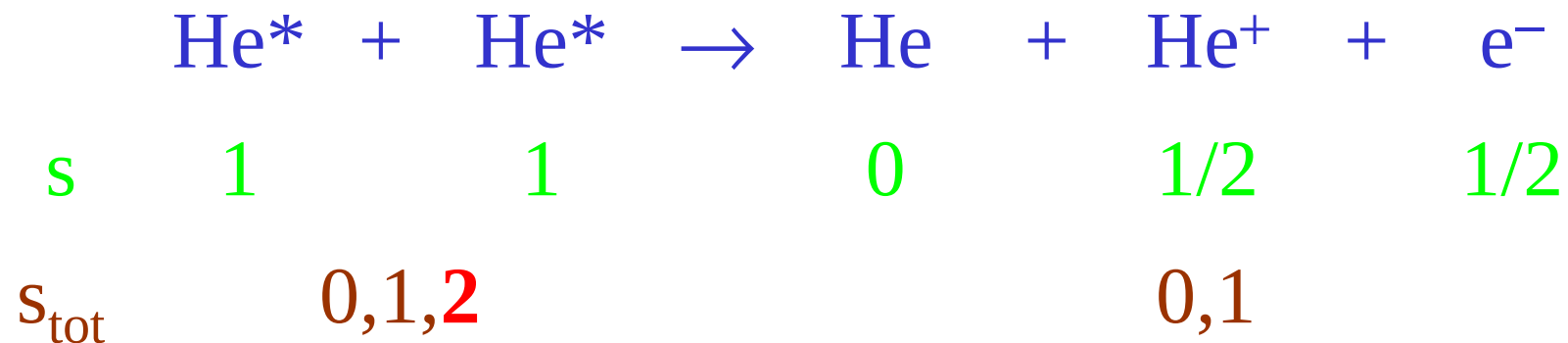
$$2^3P_2 \rightarrow 2^3S_1 : \sim 100 \text{ ns}$$

$$2^3P_2 \rightarrow 1^1S_0 : \sim 2 \text{ s}$$



Why He*?

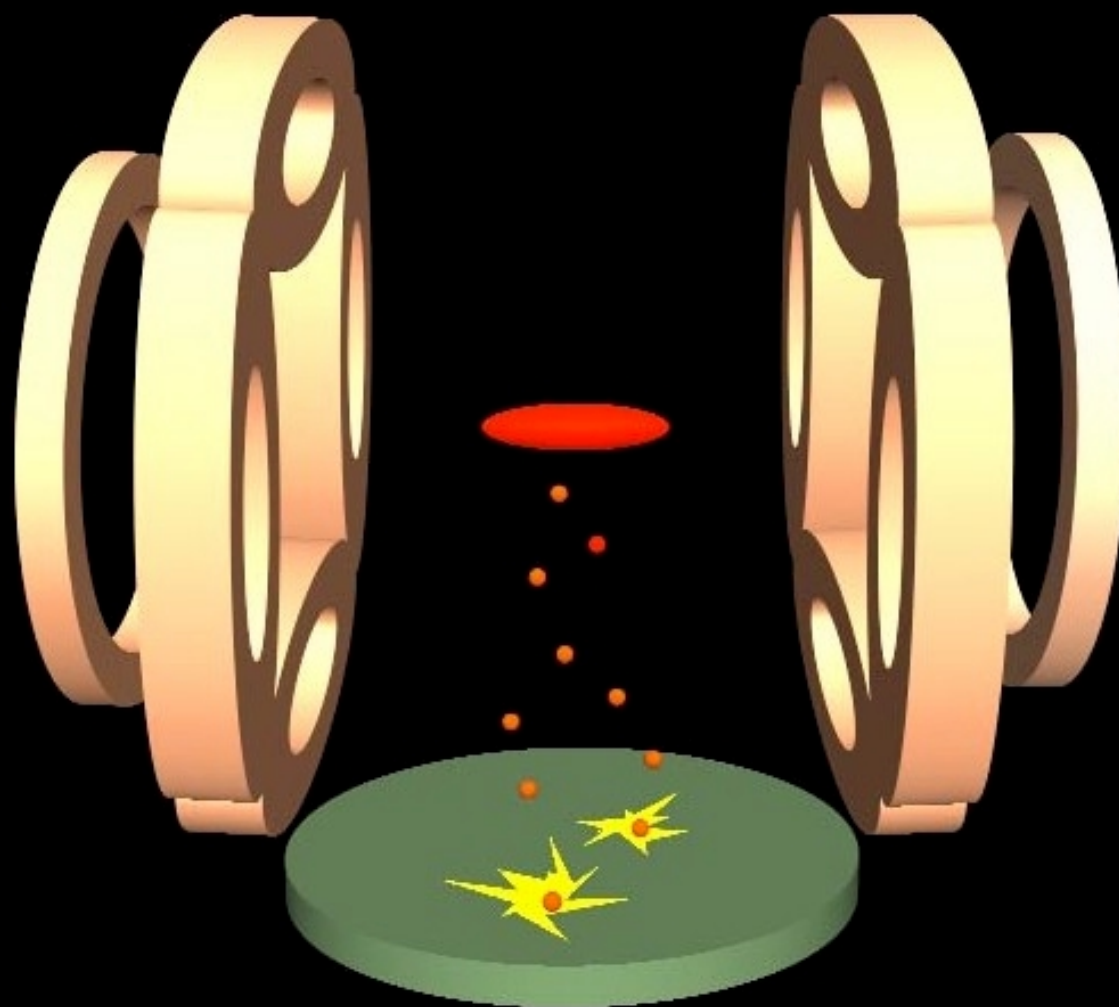
- If laser cooling of He* is a good idea, so is evaporative cooling.
- metastability: detection...
- Penning ionization



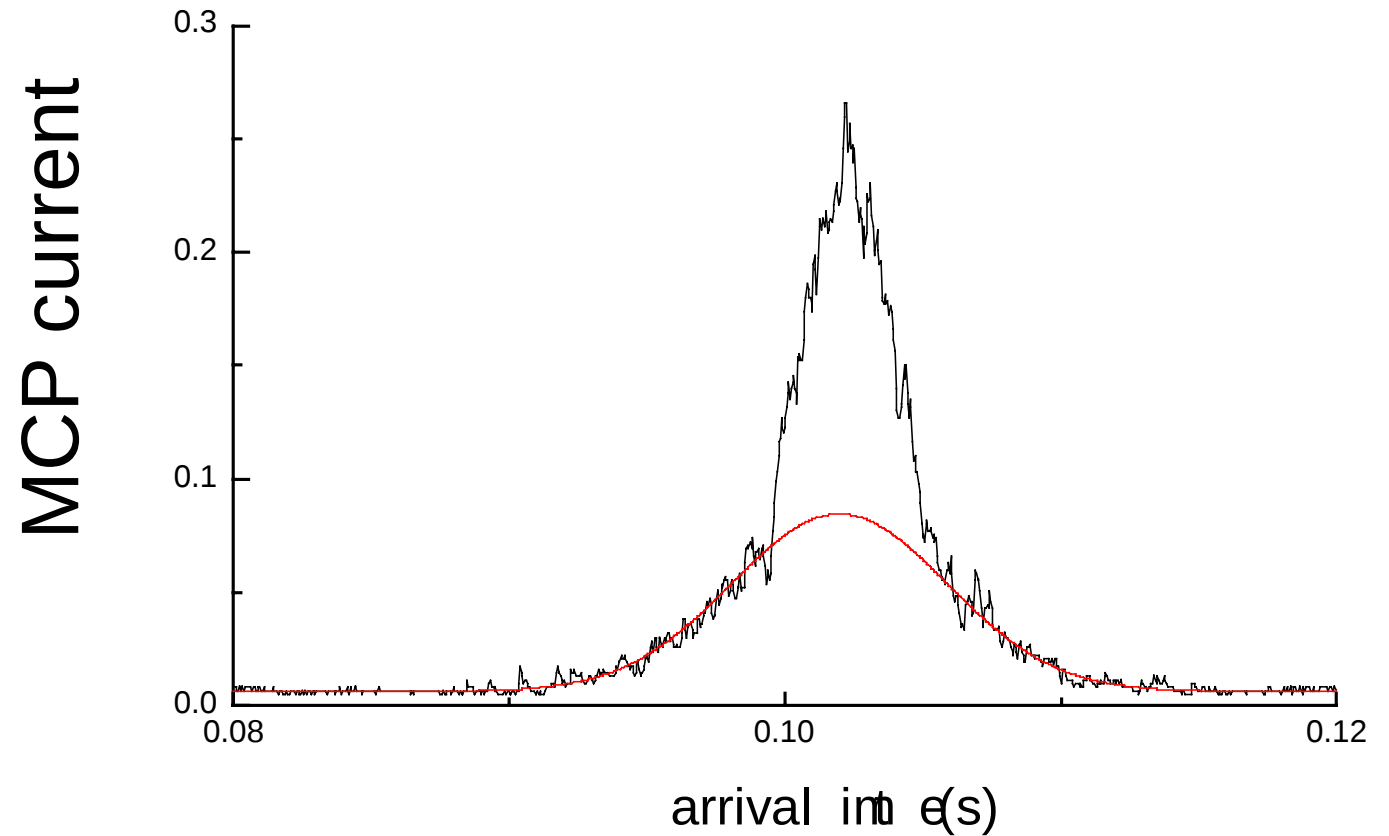
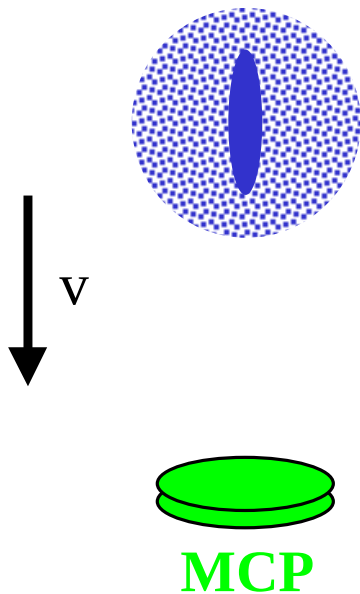
Strong suppression of ionization (Amsterdam)

Expts: Amsterdam, Utrecht, Canberra, Paris, Orsay

Apparatus



At the BEC transition



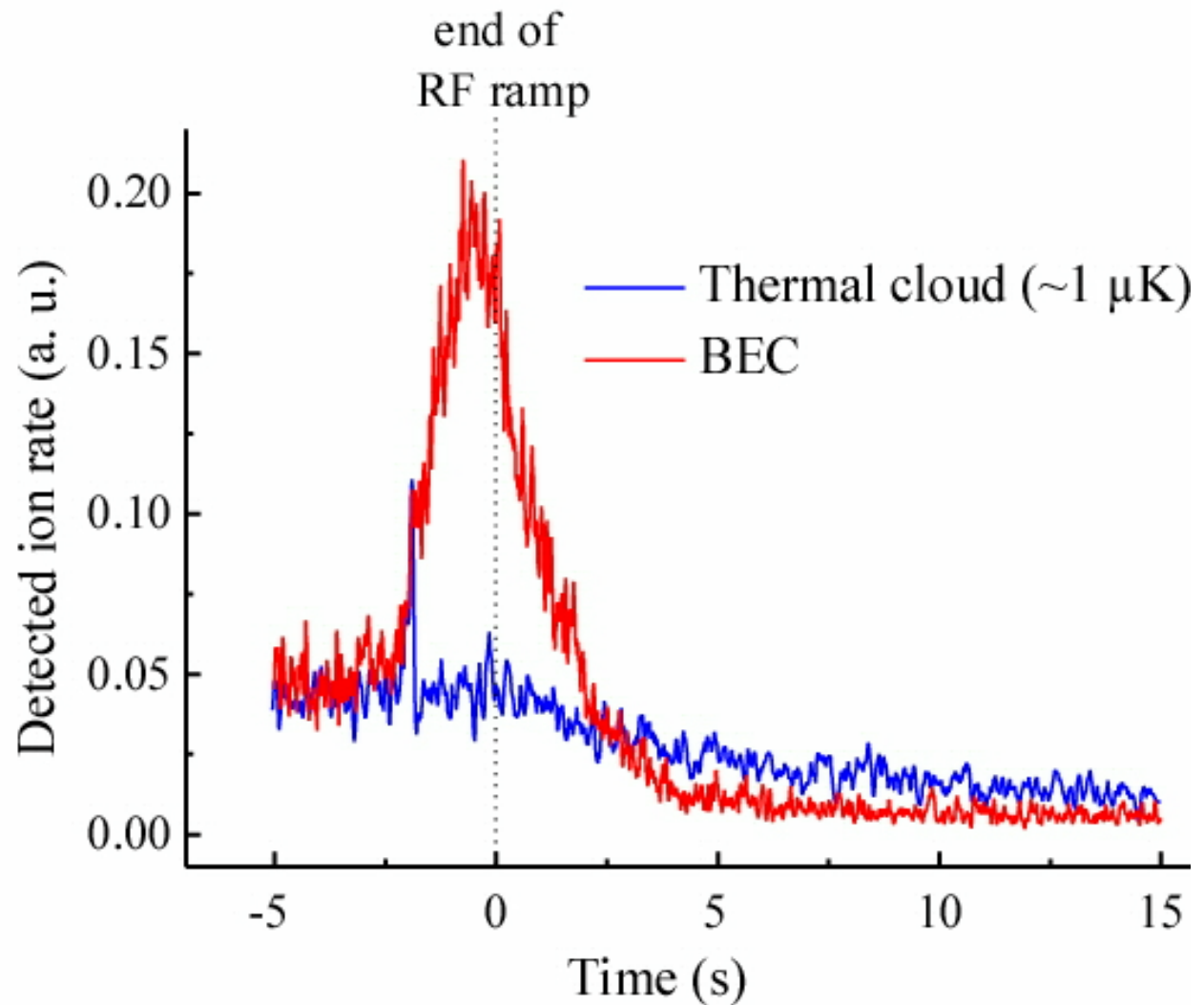
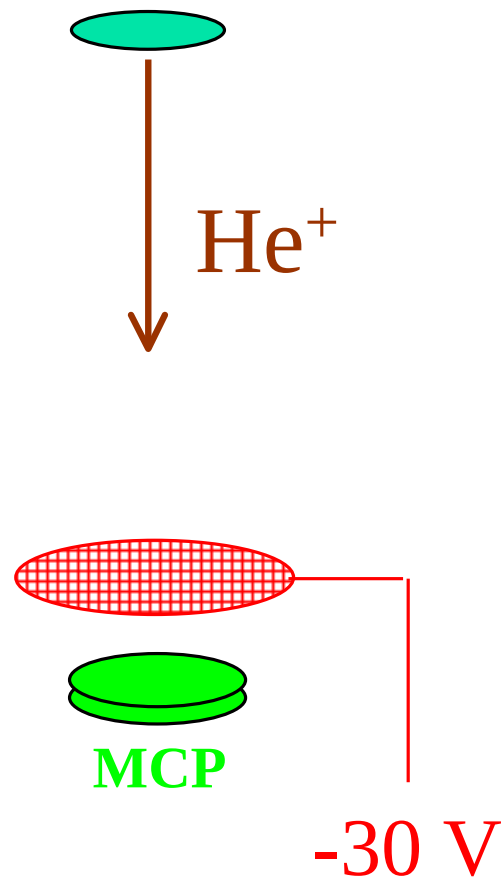
we measure N , T , μ

Summary

- $N_0 \sim 10^5$, $n_0 \sim 10^{13} \text{ cm}^{-3}$
- scattering length
 $a = 18 \pm 8 \text{ nm}$
- Thomas-Fermi parameter: $\chi = N_0 a / \sigma \sim 10^3$
- collision rate at T_C : $\gamma = n\sigma v \sim 10^3 - 10^4 \text{ s}^{-1}$
 - hydrodynamic regime
- quantum depletion $\sqrt{(na^3)} \sim 0.01$
Science, 292, 461 (2001), PRL 68, 3459 (2001)

Observation of ions

trapped atoms



loss and ionization rate

$$\frac{dN}{dt} = -\beta \int n^2 dV - L \int n^3 dV$$

$$N = f(t; \beta, L)$$

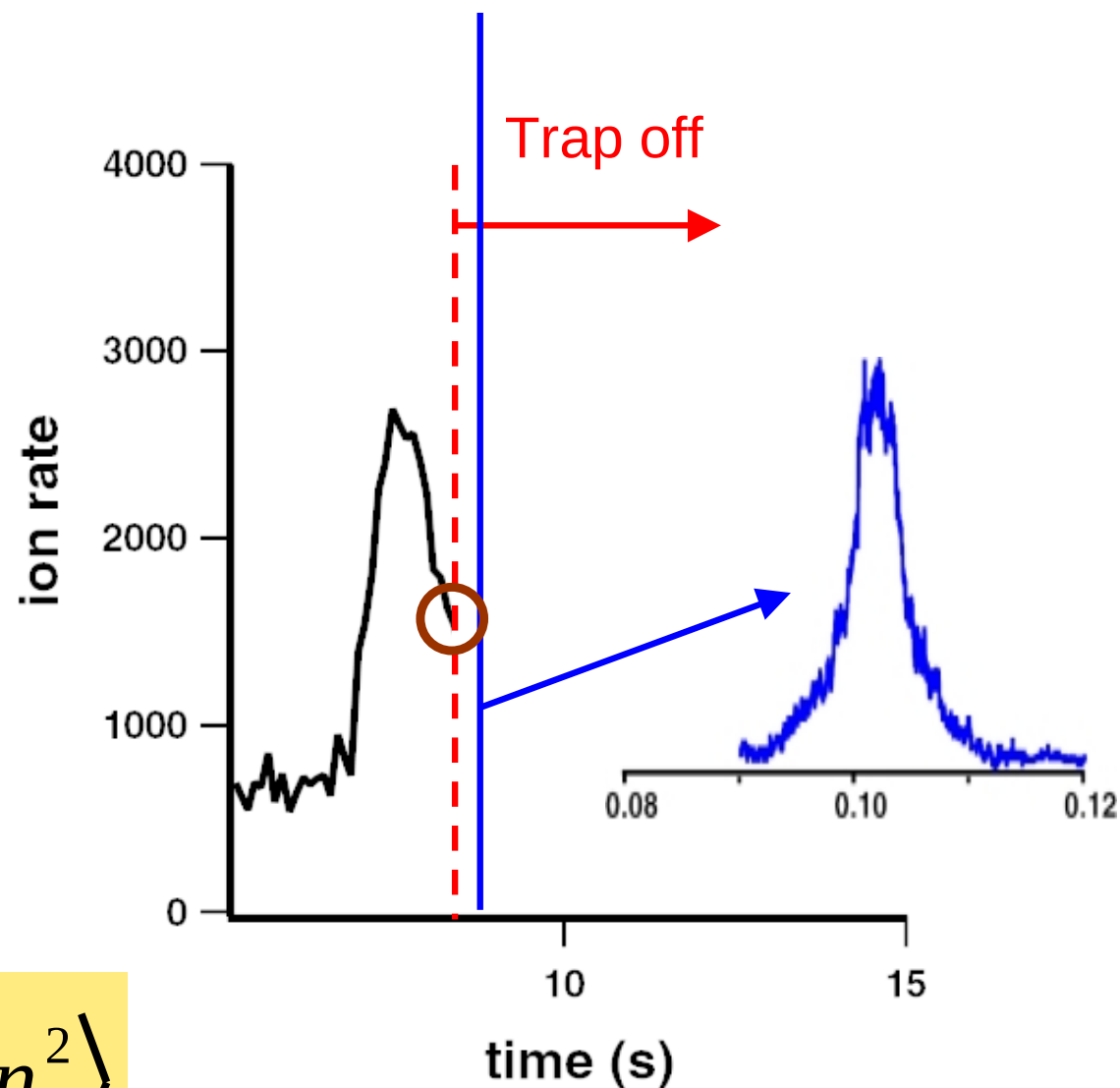
$$\frac{dI}{dt} = \frac{\beta}{2} \int n^2 dV + \frac{L}{3} \int n^3 dV$$

$$\Gamma = \frac{1}{N} \frac{dI}{dt} = \tilde{\beta} \langle n \rangle + \tilde{L} \langle n^2 \rangle$$

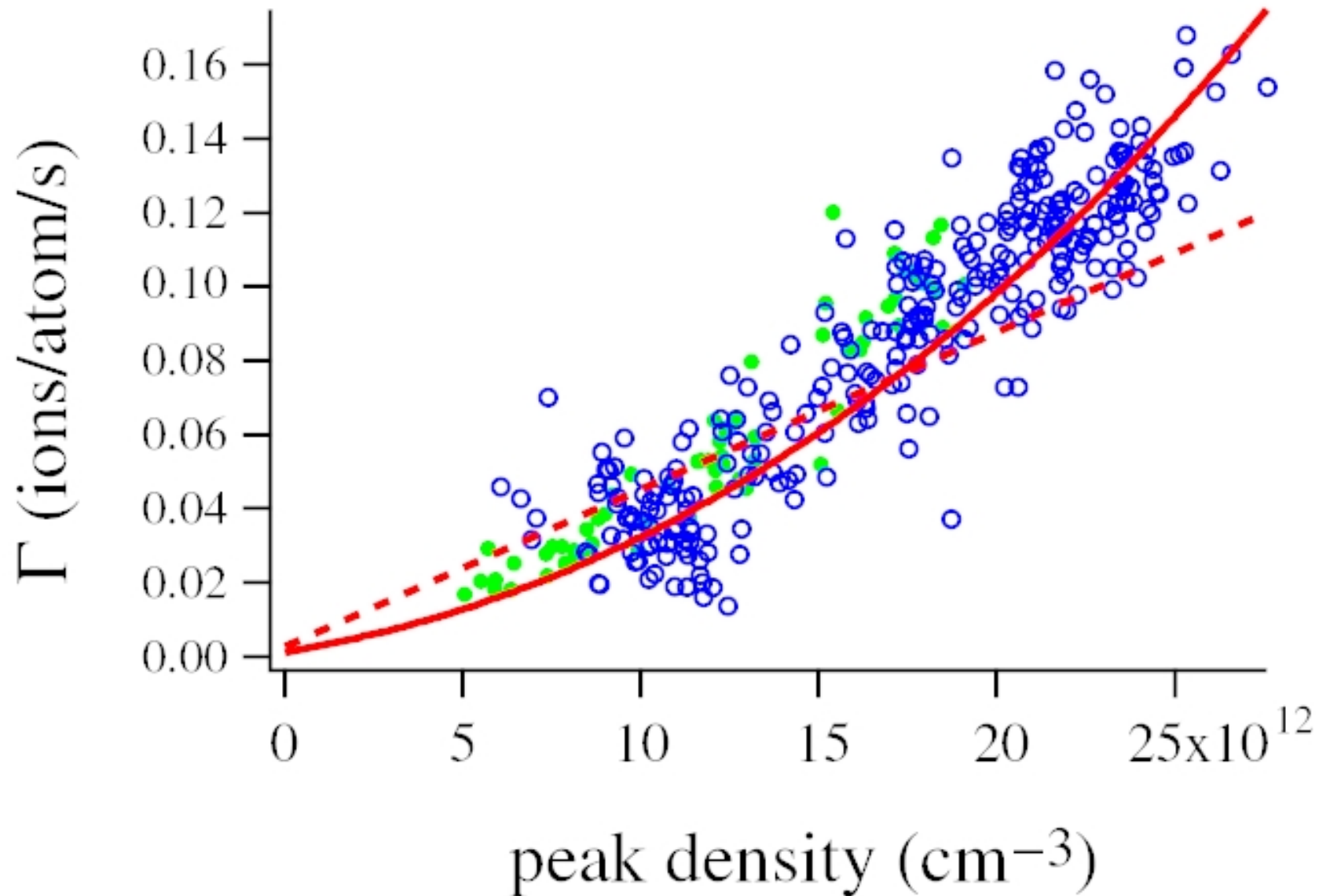
Analysis of ion signal

- lifetime: difficult to analyse
- we can measure dI/dt
- and measure $N, T, \mu \Rightarrow \langle n \rangle$ best for a "pure" BEC

$$\Gamma = \frac{1}{N} \frac{dI}{dt} = \beta \langle n \rangle + L \langle n^2 \rangle$$

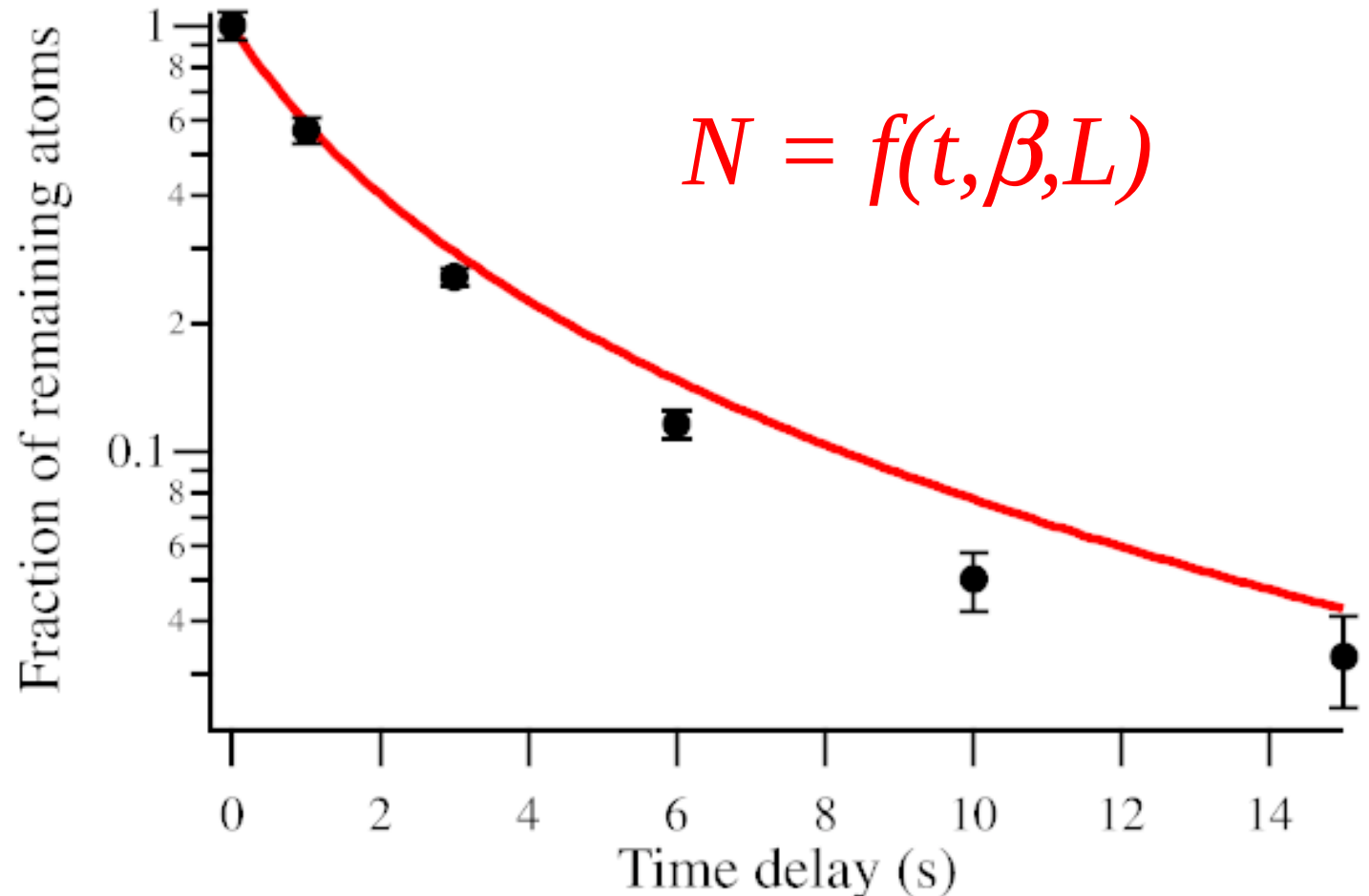


ion rate vs density (quasi-pure BEC)



decay of the number of atoms

- Independent of the ion detection
- Ionization accounts for most of the loss



conclusions

- Calibration of N and I is difficult. We use a to get N_0 .
- 2 body ionization rate coeff: $3 \times 10^{-14} \text{ cm}^3 \text{ s}^{-1}$
- 3 body ionization is important $8 \times 10^{-27} \text{ cm}^6 \text{ s}^{-1}$
- Statistical factors (2! or 3!) are important, but:

$$\Gamma_{3\text{-body}}^{\text{BEC}} = \frac{1}{3!} \left[1 + \frac{64}{\sqrt{\pi}} \sqrt{n_0 a^3} \right] L n_0^2$$

Kagan, Svistunov, Shlyapnikov, **50% correction**
JETP Lett. 42, 209 (1985)

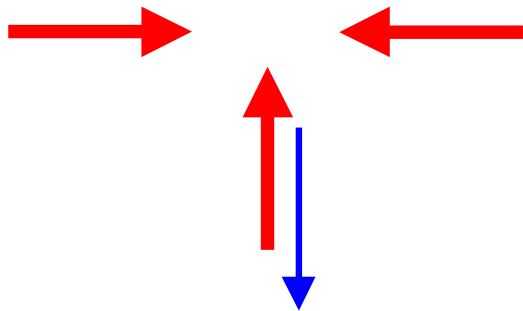
- We need the same for a thermal cloud
- and a better measurement of a .

making correlated atom pairs

Four wave mixing

NIST: Science 398, 218 (1999)

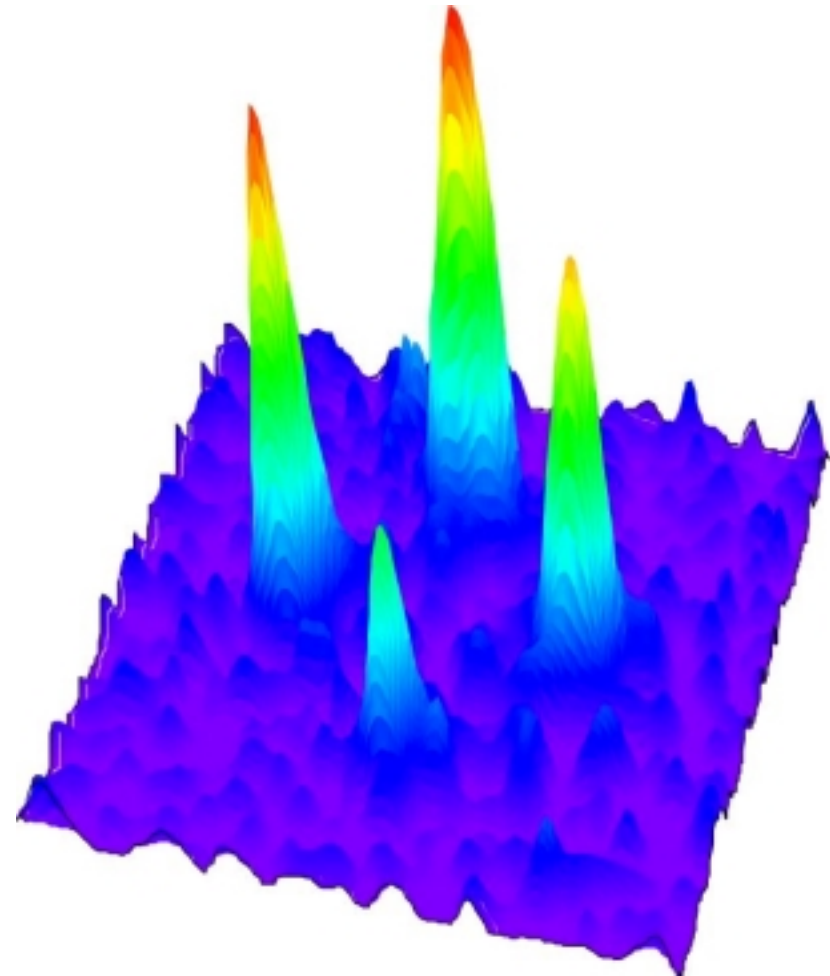
MIT: cond-mat/0203286 (2002)



Atoms (photons) created
in entangled pairs

PRL 85, 3987 (2000)

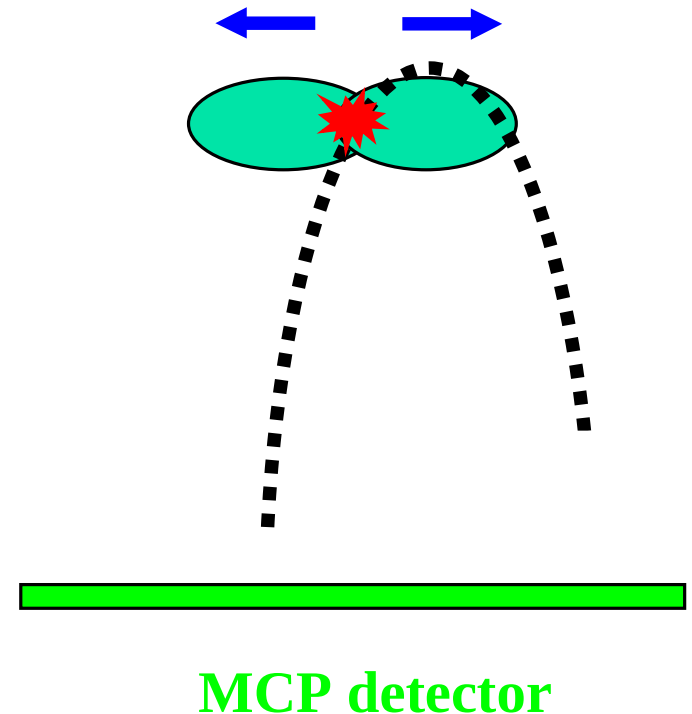
PRL 85, 3991 (2000)



Data from NIST 1999

detecting the pairs

- Spontaneous 4 wave mixing ("perfect correlations")
- Raman pulses create colliding, untrapped condensates (other ways?)
- Spatial correlations become temporal correlations
- $|0\rangle + |0\rangle$ collisions: Inelastic rate !



Thank you

- groupe 'hélium':

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